



Lecture 20

Magnetism, Magnetic Order, and Superconductivity

Love thy neighbors

Origin of Magnetic Moment in Solids

- Electrons are the main source
(Nuclear spins do come in handy as in NMR)
- Magnetic moment
 - $\boldsymbol{\mu} = -\mu_B (\mathbf{L} + g_0 \mathbf{S}), g_0 = 2.0023\dots \approx 2$
 - $\mu_B = e\hbar/(2m)$
 - $\mu_B = 9.27e-24$ J/T (SI) or $5.8e-9$ eV/gauss
 - For nucleons, reduction by ~ 2000
 - Remember, B field ~ 100 T = $1e6$ gauss is the strongest field that one can generate, so magnetic perturbation energies are quite small.

Some Atomic Physics ...

- **Spin-orbit coupling:**

- $H_{LS} = \mathbf{L} \cdot \mathbf{S} \frac{e^2}{r^3 8\pi\epsilon_0 m^2 c^2}$
- Goes like $\sim \alpha^2 \times$ unperturbed energy
- The heavier the atom, the greater $\langle H_{LS} \rangle$
- The closer the orbital is to the atom, the greater $\langle H_{LS} \rangle$
- Quite important for the band structure of semiconductors (44 meV for Si, 0.34 eV for GaAs)

- **Zeeman interaction:**

- $H_B = -\boldsymbol{\mu} \cdot \mathbf{B} = -\mu_B (\mathbf{L} + 2\mathbf{S}) \cdot \mathbf{B}$

- For very light atoms (Li, e.g.), the Zeeman term can be more important than the spin-orbit coupling.

- **For most solid problems, the spin-orbit coupling is more important than the Zeeman term.**

- Treat spin-orbit first, and then, treat Zeeman as a perturbation.
- Good Quantum Numbers: J, L, S, J_z
- $\langle H_B \rangle = g(L, S, J) \mu_B J_z B$ (**g: Lande g-factor; eq. 7.10; see Sakurai**)

$$\mu_{\text{eff}} = g \mu_B J_z$$

Hund's Rule for Atomic Spectra

- Empirical rule that works pretty well for 3d Transition metal (TM) and 4f rare earth (RE) ions
- Many electrons in an un-filled orbital
 - First rule: maximize S
 - Second rule: maximize L
 - Third rule: $J = |L-S|$ if less than half full and $J = L+S$ if more than half full
 - Notation: $^{2s+1}L_J$
- E.g., $V^{3+} : 3d^2 \ ^3F_2$, $Fe^{2+} : 3d^6 \ ^5D_4$, $Mn^{2+} : 3d^5 \ ^6S_{5/2}$
- Roughly, the physics is due to Coulomb interaction and spin-orbit interaction

Curie Paramagnetism

- Localized magnetic moments in response to external B field
- $M = \chi H$
- Curie law at high T, small B: $\chi = C/T$
 - $C = N p^2 \mu_B^2 \mu_0 / (3 k_B)$ (N is volume density)
 - $p = g [J(J+1)]^{1/2}$
- Saturation at low T, large B: $M = Ng\mu_B$
- p agrees well for RE, but not so for TM
- For TM ions, $p = 2 [S(S+1)]^{1/2}$ works much better!?

Quenching of Orbital Moment in Transition Metal Ions in Crystal

- Empirically observed (Kittel, p. 308)
- Effect of Crystal Field (large ~ 1 eV in TM ions)
- L_z is not a good quantum number
- $\langle L_z \rangle$ (also for x,y) is often zero

Pauli Para-magnetism of Metals

- Response of Fermi sea to the magnetic field
- Spin-up electron's energy goes up and spin-down ... down, both by $\mu_B B$
- $M = \mu_B(n_{\uparrow} - n_{\downarrow}) = \mu_B^2 g(E_F) B$ (g is per volume, but for both spins)
- $\chi = M / H \approx \mu_0 \mu_B^2 g(E_F)$ (T independent)
- $\chi \sim \mu_0 \mu_B^2 N / E_F$ (N is volume density), i.e. weakened by T/T_F (quantum behavior – only those at E_F respond) compared to Curie susceptibility (classical behavior – all spins respond)

Diamagnetism

- Lenz's law
- Landau Diamagnetism
 - Orbital Motion of Electrons
 - For free electron, $\chi_{\text{Landau}} = -\chi_{\text{Pauli}}/3$
 - For effective mass m^* , $\chi_{\text{Landau}} / \chi_{\text{Pauli}} \sim (m/m^*)^2$
i.e. $|\chi_{\text{Landau}}| \gg \chi_{\text{Pauli}}$ (semi-cond) and $|\chi_{\text{Landau}}| \ll \chi_{\text{Pauli}}$ (heavy Fermion)
- Larmor (Langevin) Diamagnetism due to closed shell ions ($\mathbf{A} = -\mathbf{r} \times \mathbf{B} / 2$, $H = (\mathbf{p} + e\mathbf{A})^2 / 2m$, The last term of $H \sim r^2 B^2$)
- For non-magnetic materials, often it is necessary to consider all susceptibilities (Pauli, Landau, Larmor, and Curie – often due to magnetic impurities)

Different Kinds of Magnetic Order

- **Ferro-magnetism** (Fe, CrO_2 , EuO , etc.)
- **Anti-ferromagnetism** (MnO , FeF_2 , HTSC cuprates, etc.)
- **Ferri-magnetism** (Fe_3O_4) ...
- Note
 1. Magnetic orders happen at fairly high temperatures (100 K ~ 1000 K) – i.e. internal field is very high due to Coulomb interaction
 2. The above are $B=0$ properties, while the following designate response to finite B :
 - Para-magnetism (Curie, Van Vleck, Pauli)
 - Dia-magnetism (Larmor, Landau, London)

Interaction between Magnetic Ions

- Dipole-dipole interaction is very small

$$\vec{B} = \frac{3(\vec{m} \cdot \vec{r})\vec{r} - \vec{m}}{r^3} \quad (\text{cgs})$$

$$\mu_B \approx 6 \times 10^{-9} \text{ eV/gauss} \approx 9 \times 10^{-21} \text{ erg/gauss}$$

$$B \sim \frac{\mu_B}{r^3} \sim \frac{9 \times 10^{-21}}{(3 \text{ \AA})^3} \sim \frac{10^{-21}}{3 \times 10^{-24}} \sim 1000 \text{ gauss}$$

$$E \sim \mu_B B \sim 6 \times 10^{-6} \text{ eV} \sim 0.006 \text{ meV} \sim 0(0.01) \text{ K}$$

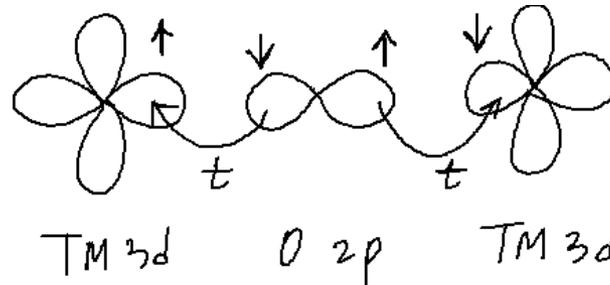
- Magnetic exchange energy results from Coulomb interactions: hopping (t) and exchange interactions
- Two most important Hamiltonians:

Heisenberg $H = -\sum_{\langle i,j \rangle} J_{ij} \vec{S}_i \cdot \vec{S}_j$

Hubbard $H = \text{Kinetic Energy} + \text{Exchange Coulomb Energy}$

Sign of J in Heisenberg Model

- Positive like “Hund’s rule” ($J > 0$): Equal spins repel less
- **Super-Exchange** ($J < 0$): TM ions (such as ions of Fe, Mn, Cu etc) interact via anions



- Double exchange ($J > 0$), RKKY (J can be random), ...

Ferromagnetism due to local spins

- Consider Heisenberg Hamiltonian with only nearest-neighbor interaction (now $\langle i,j \rangle$ means only nearest neighbor pairs, not all possible pairs) and $J > 0$

$$H = -J \sum_{\langle i,j \rangle} \vec{S}_i \cdot \vec{S}_j$$

- This Hamiltonian is very difficult to solve. One can however show that all spins pointing in the same direction is the lowest energy state. For instance for spin $1/2$, it is easy to show

$$H |\psi\rangle = -NJ/4 |\psi\rangle \quad |\psi\rangle = |\uparrow\uparrow\uparrow \dots \uparrow\uparrow\uparrow\rangle$$

where z direction is any random direction. (Hint: use $S_x = \frac{1}{2} (S_+ + S_-)$ $S_y = \frac{-i}{2} S_+ + \frac{i}{2} S_-$)

Ferromagnetism due to local spins

- Clearly the ground state is one in which all spins are aligned (symmetry breaking) and the excited state is in one in which spins are flipped (magnon or spin wave)
- At finite T , each spin will “fluctuate” around its mean value (even at $T=0$ due to quantum fluctuations)
- An elementary approach is then to ignore these fluctuations and apply a “mean field” theory

Mean field theory of FM (Weiss)

- There is an enormous internal field due to magnetization ($\lambda \gg 1$):

$$\lambda \mu_0 \vec{M} \quad \vec{B} = \lambda \mu_0 \vec{M} + \vec{B}_{loc}$$

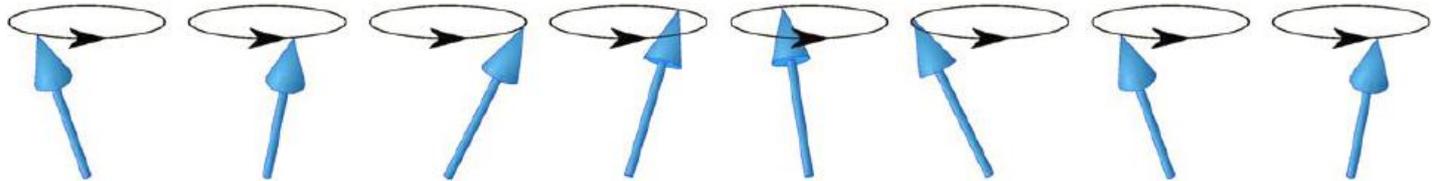
- At high T, Curie-Weiss Law: $\chi = \frac{C}{T - T_c}$
- At $T < T_c$, spontaneous magnetization
- $T = 0$: Saturation without any external field $M = N\mu$
- Provides basic picture for the FM transition but ...
 - T-dependence near T_c (wrong):

$$M \propto \left(1 - \frac{T}{T_c}\right)^{\frac{1}{2}} \quad \chi \propto (T - T_c)^{-1}$$

- T-dependence near $T = 0$ (wrong):

$$M = N\mu \left(1 - 2 \exp\left(-\frac{2T_c}{T}\right)\right)$$

Spin Wave and Bloch $T^{3/2}$ law in FM



<http://upload.wikimedia.org/wikipedia/commons/0/01/FerromagneticMagnon.svg>

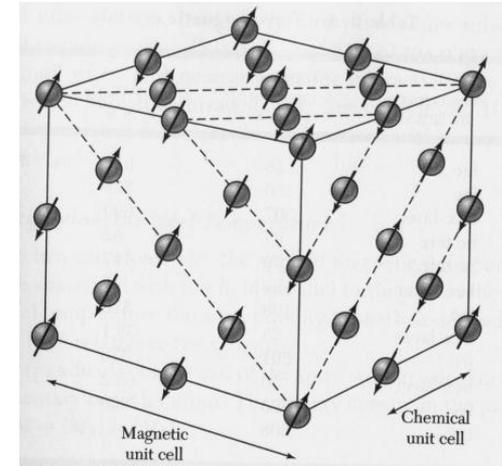
- Quantum Mechanically, spin wave means excited states with one spin flipped
- Easiest way to consider this is to do Classical Mechanics (like phonon problem)
- We know that $k=0$ solution must exist with energy = 0 (like the acoustic phonon case)
- FM spin wave has the property: $\omega \propto k^2$
- This leads to Bloch's law near $T=0$:

$$C \propto T^{3/2} \quad \vec{M} \text{ reduction} \sim T^{3/2}$$

Antiferromagnetism

- Spins align anti-parallel to each other ($J < 0$ due mostly to the Anderson super-exchange)
- Different from FM in that
 - AFM state is not an eigenstate
 - Not even ground state (!)
 - But thermodynamically favored state at $T < T_N$
 - Magnetic unit cell differs from chemical unit cell
 - Spin wave is really like phonon:

$$\omega \approx \omega_{ex} |ka| \quad k \rightarrow 0$$



Mn²⁺ ions in MnO (from Kittel)

High T susceptibility of FM or AFM due to local spins

$$\chi \sim 1/(T - \Theta)$$

(FM: $\Theta \sim O(T_C)$, Curie Temperature)

$$\chi \sim 1/(T + \Theta)$$

(AFM: $\Theta \sim O(T_N)$, Néel Temperature)

Θ is not exactly T_C (or T_N) because the theory presented here is very simple (only nearest neighbor interaction and mean-field)

Itinerant Ferromagnetism

- One of the mechanism for metallic FM (as in Fe)
- Due to exchange Coulomb interaction
- Exchange interaction favors spin polarization but kinetic energy favors no spin polarization (**competition**)
- The onset of the FM is given by Stoner criterion in perturbation MF theory

$$\frac{g(E_F)}{N} \cdot U = 1$$

Stoner Ferromagnetism

- Just like Pauli Para-magnetism theory, but with internal field

$$E_{\vec{k}\uparrow} = E_{\vec{k}} + U n_{\downarrow} \quad \text{Exchange energy } U > 0$$

$$E_{\vec{k}\downarrow} = E_{\vec{k}} + U n_{\uparrow}$$

$$M = \mu_B (n_{\uparrow} - n_{\downarrow})$$

$$E_{\vec{k}\uparrow} = E_{\vec{k}} + U \cdot \frac{n}{2} - U \frac{M}{2\mu_B}$$

$$n = n_{\uparrow} + n_{\downarrow}$$

$n, n_{\uparrow}, n_{\downarrow}$: # densities

$$E_{\vec{k}\downarrow} = E_{\vec{k}} + U \cdot \frac{n}{2} + U \frac{M}{2\mu_B}$$

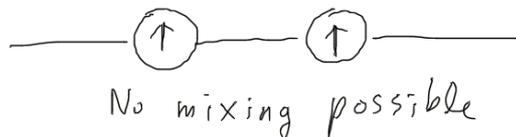
- Self-consistency (source M is the same as the result, as in Weiss MF theory) – left for reader – leads to “Stoner Criterion”

$$\frac{g(E_F)}{N} \cdot U = 1$$

Hubbard Model and AFM

- Hubbard Model provides a basis for the Stoner FM (small U)
- Hubbard Model also provides a basis for an AFM in the opposite limit of large U

Hubbard model has an AFM when
 $U \rightarrow \infty$



Also, Mott-Hubbard Insulator



I.e., AFM MH Insulator

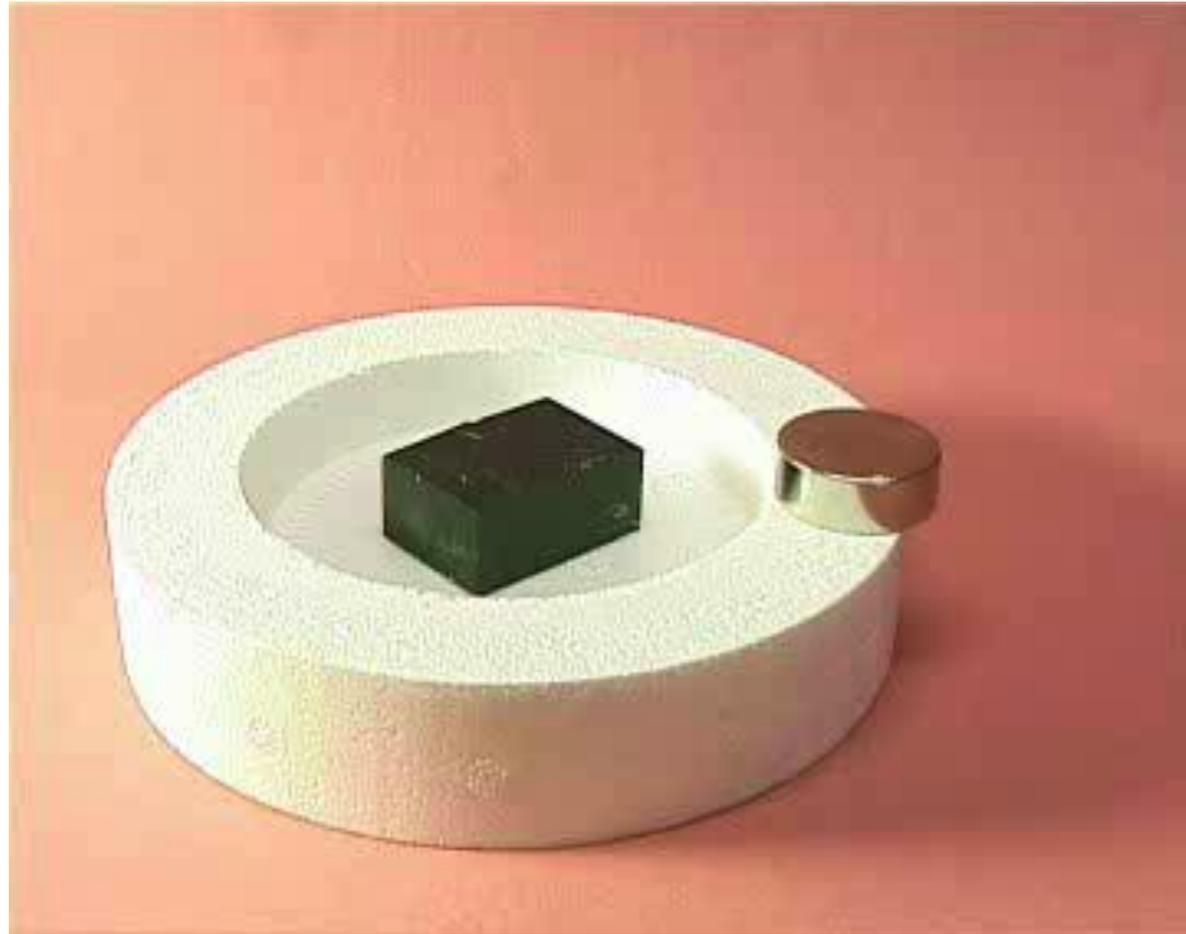
E lowered by $-\frac{t^2}{U}$



Superconductivity

- Perfect diamagnetism
- Perfect conductivity

Meissner Effect



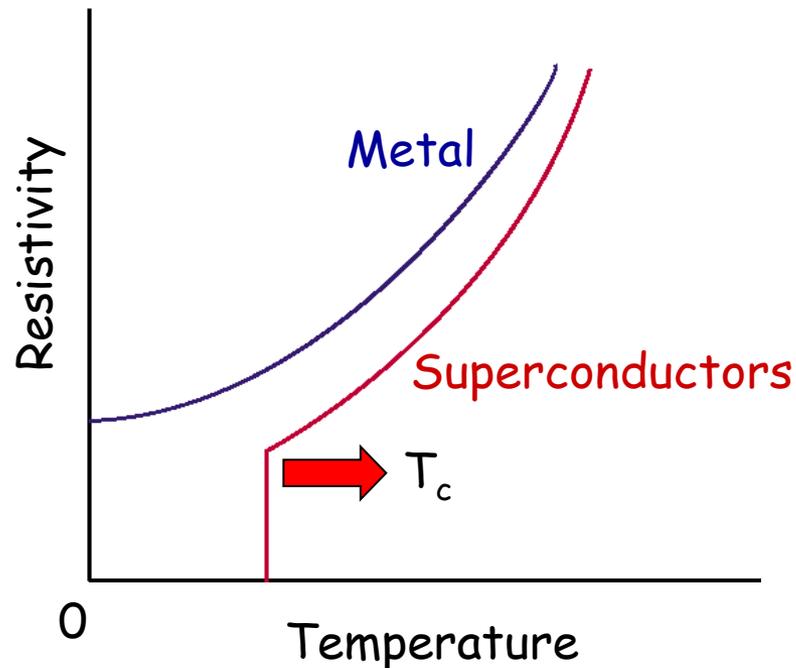
<http://www.fys.uio.no/super/levitation/>

Essential Characteristics

1911 K. Onnes Superconductivity in Hg

1933 Meissner effect

RESISTANCELESS CONDUCTION



MEISSNER EFFECT:
Perfect diamagnetism

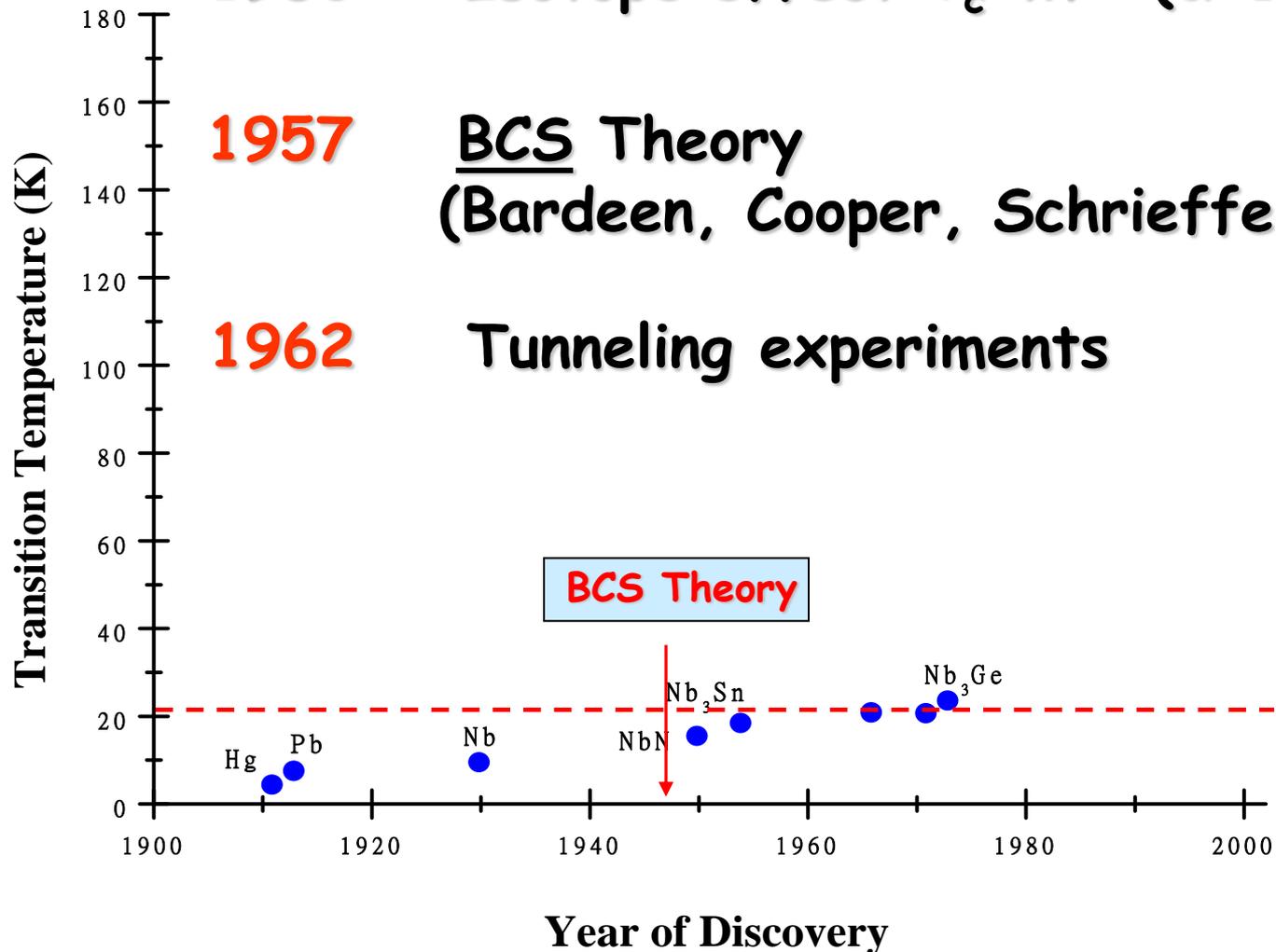


Understanding of Superconductivity

1950 Isotope effect $T_c \sim M^{-\alpha}$ ($\alpha \sim 1/2$)

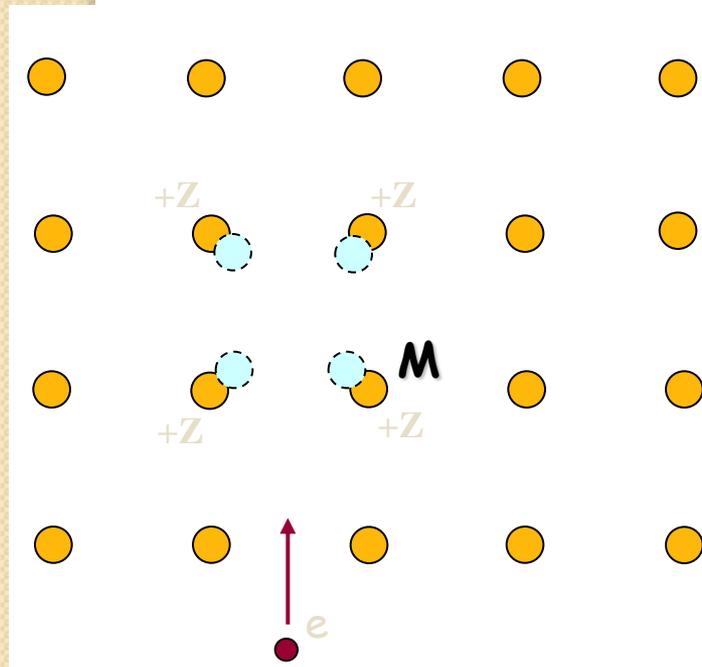
1957 BCS Theory
(Bardeen, Cooper, Schrieffer)

1962 Tunneling experiments



BCS theory

PHONON MEDIATED PAIRING (phonon = lattice vibration)



Pairs of electrons: **Cooper pairs**

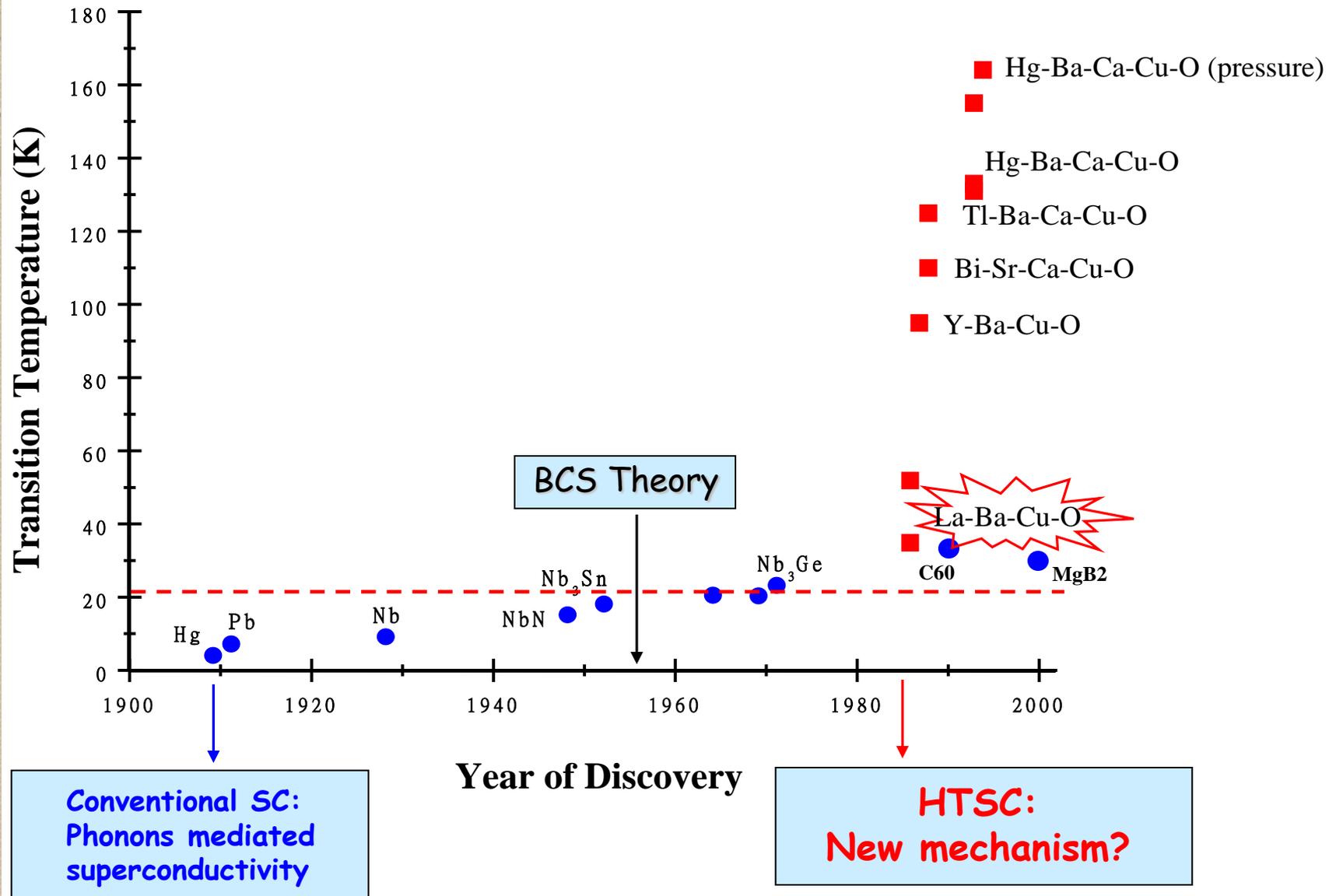
Superconducting gap: Δ

EI-ph coupling constant: λ

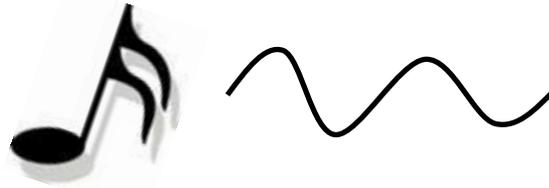
$$T_c \sim \omega_{ph} \exp(-1/\lambda) \sim M^{-1/2}$$

e-ph wins e-e at low freq. (Slow Wins!)

HTSC – New Superconductivity



Superconductivity – Dance of Electron Pairs



Origin of SC = **PAIRS** dancing to the **SAME** tune



So How does it work?

$$\vec{j}_0 = \frac{\hbar q}{2im} (\psi_0^* \vec{\nabla} \psi_0 - (\vec{\nabla} \psi_0)^* \psi_0)$$

QM, e.g. Sakurai

$$\vec{j} = \vec{j}_0(\psi_0 \rightarrow \psi) - \frac{q^2 \vec{A}}{mc} |\psi|^2$$

$$H = \frac{1}{2m} \left(\vec{p} - \frac{q\vec{A}}{c} \right)^2 + \text{pot.}$$

$\psi \approx \psi_0$ (energy gap, many-body coherence)

$$\vec{j}_0 = 0$$

$$\vec{j} = -\frac{q^2 \vec{A}}{mc} n_s$$

London Equation

$$\vec{\nabla} \times \vec{B} = \frac{4\pi}{c} \vec{j}$$

$$q = -2e, \quad m \approx 2m_e$$

Cooper Pair

$$-\vec{\nabla}^2 \vec{A} = \frac{4\pi}{c} \vec{j} = -\frac{4\pi q^2 n_s}{mc^2} \vec{A}$$

$\lambda \sim$ a few 100 Å

$$\vec{\nabla}^2 \vec{A} = \frac{\vec{A}}{\lambda^2}$$

$$\lambda = \sqrt{\frac{mc^2}{4\pi q^2 n_s}}$$

\vec{B} field is screened within the length scale λ

Meissner Effect

Steady state No electrostat. pot.

Infinite Conductivity

$$\frac{\partial \vec{j}}{\partial t} = 0$$

$$\vec{E} = -\frac{1}{c} \frac{\partial \vec{A}}{\partial t}$$

$$\vec{E} = 0$$

Two Length Scales and two types of SC

λ = London Penetration Depth

ξ = coherence length $\sim \frac{\hbar v_F}{\Delta}$

Δ : pair binding energy

$\xi \sim$ Pair wavefunction size

Pippard non-local E & M \leftarrow due to ξ

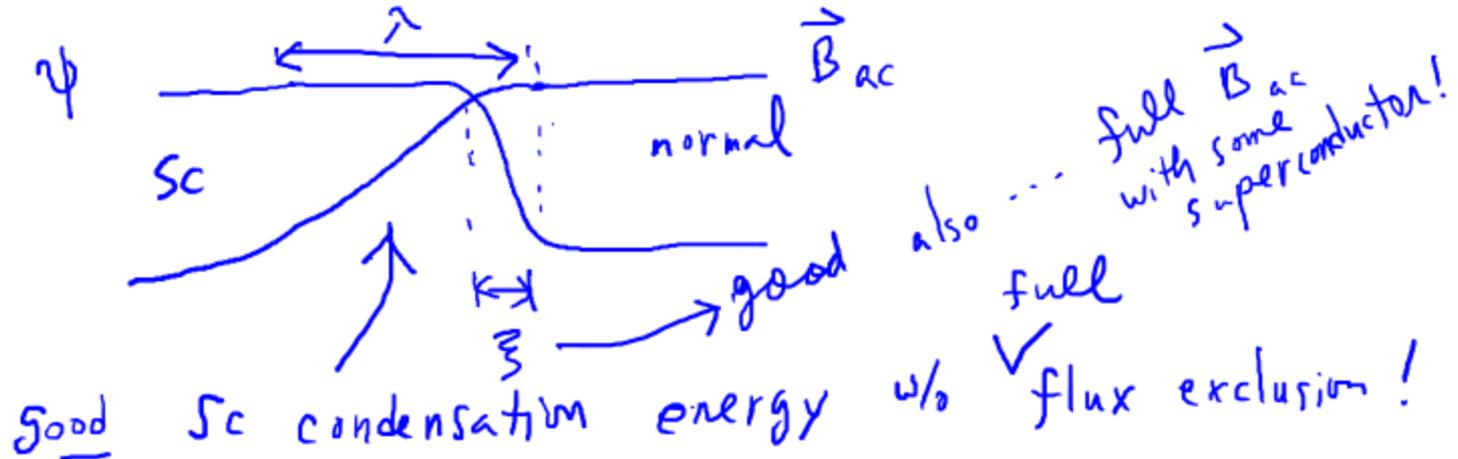
As material becomes impure, λ and ξ change in opposite manners.

Type 1 SC: $\lambda(T=0) \ll \xi$

Type 2 SC: $\lambda(T=0) \gg \xi$ (most useful)

Two Length Scales and two types of SC

Type II



Type I

